

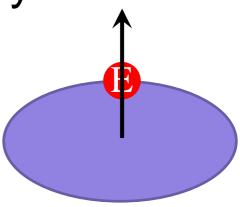
ICQM

International Center for Quantum Materials

Magnetizations, Thermal Hall Effects and Phonon Hall Effect

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Contents

- Experiment evidences of thermal Hall effect
- Issues of existing theories
- Magnetization correction to Kubo formulas
- Theory of phonon Hall effect
- Summary

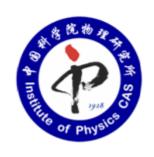
Tao Qin, Qian Niu and Juren Shi, *Energy magnetization and thermal Hall effect* Phys. Rev. Lett. **107**, 236601(2011)

Tao Qin and Junren Shi, *Berry curvature and phonon Hall effect*, arXiv: 1111. 1322 (2011)

Collaborators



Tao Qin

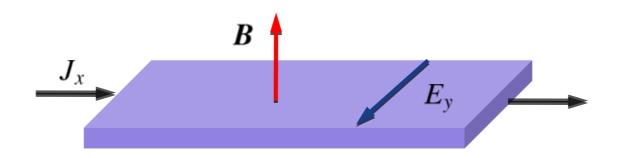




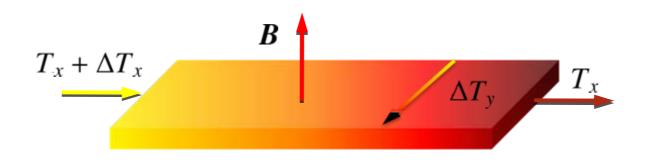
Qian Niu



Thermal Hall effect

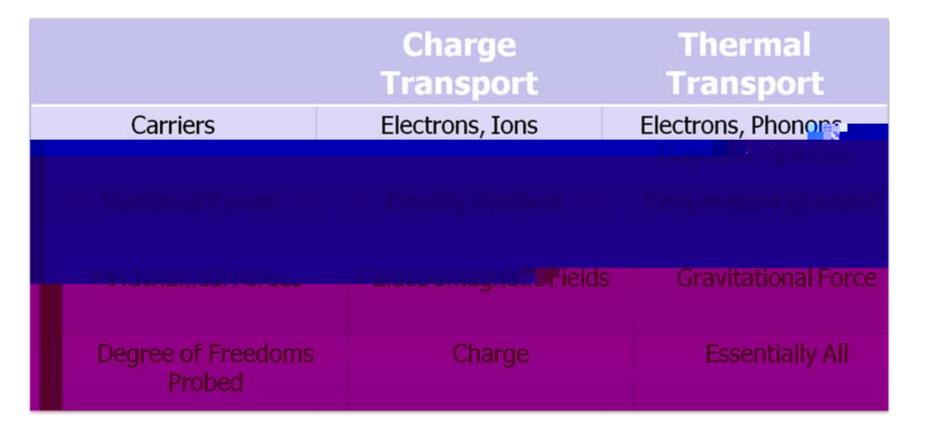


Charge Hall effect



Thermal Hall effect

Why Thermal Transport?



Thermal transport -- the more effective ways for probing condensed matter systems!

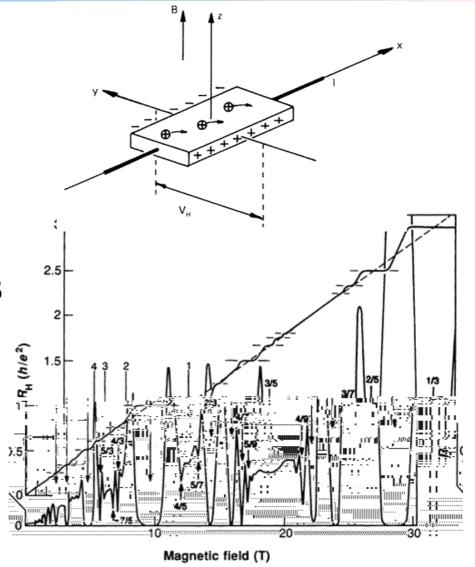
Why Hall Effect?

Quantum Hall Effect

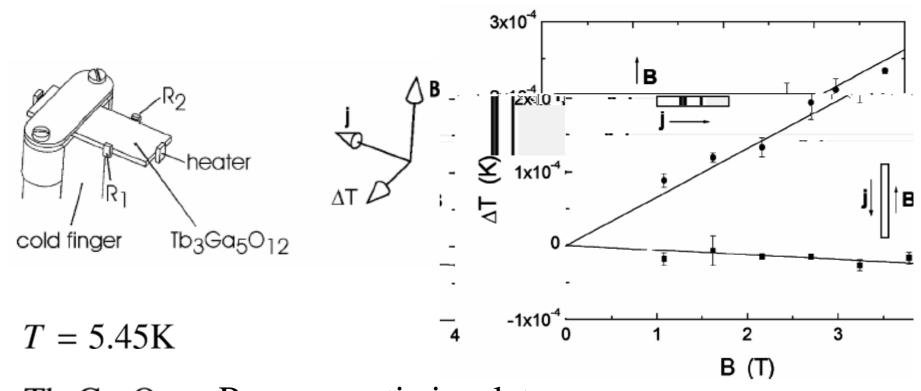


- Klitzing, 1985
- Laughlin, Stormer, Tsui, 1998

- Hall effect for the heat flow?
- Quantum Hall effect for the heat flow?



Phonon Hall effect

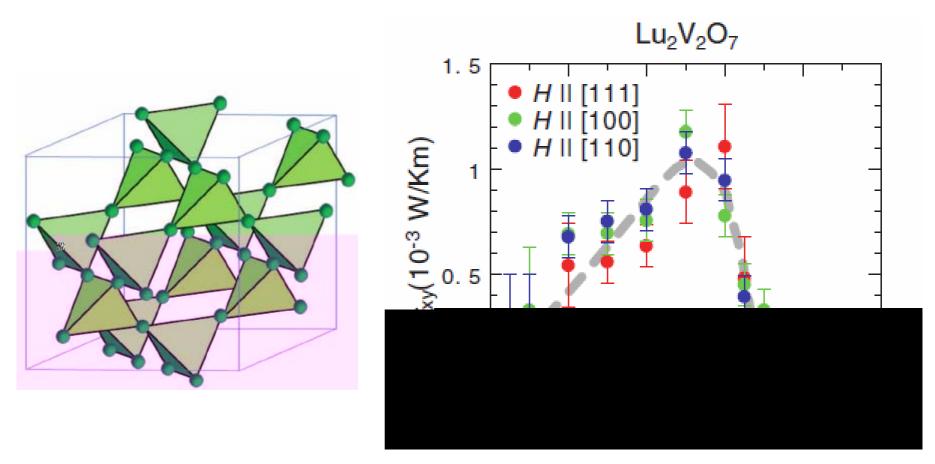


Tb₃Ga₅O₁₂ Paramagnetic insulator

C. Strohm et al., Phys. Rev. Lett. 95, 155901(2006).

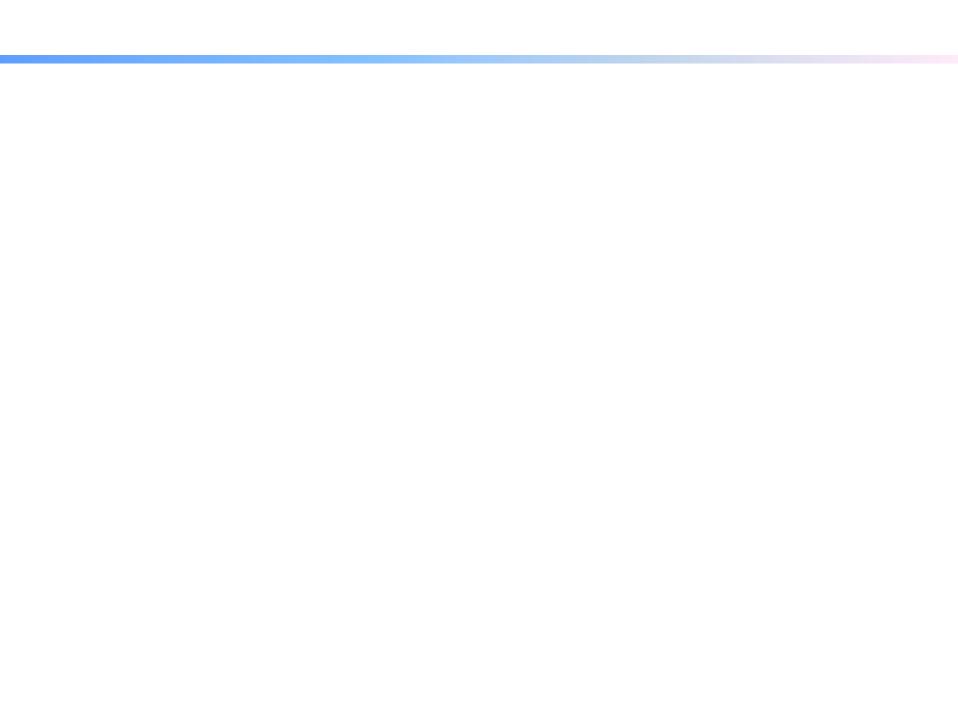
A.V. Inyushkin et al., JETP Lett. 86, 379 (2007).

Magnon Hall effect



Lu₂V₂O₇ Insulating collinear ferromagnet

Y. Onose, et al., Science 329, 297 (2010).



Kubo Formula

Thermal Hall Coefficient:
$$\kappa_{xy} = \frac{J_{Qx}}{\partial_y T}$$

Standard tool for evaluating transport coefficients:

$$\kappa_{xy} = \frac{1}{T} \int_0^\infty dt e^{-st} \beta \left\langle \hat{J}_{Qy}(0); \hat{J}_{Qx}(t) \right\rangle$$

$$\langle \hat{q} : \hat{h} \rangle = \frac{1}{\sqrt{2}} \int_{0}^{\beta} d\lambda \operatorname{Tr} \left\{ \hat{q}_{1} \operatorname{eva}(\lambda \hat{H}) \hat{q}_{1} \operatorname{eva}(\pi, \lambda \hat{H}) \hat{h}_{1} \right\}$$

Mahan, *Many Particle Physics*Kubo, Toda and Hashitsume, *Statistical Physics II*

Kubo formula applicable?

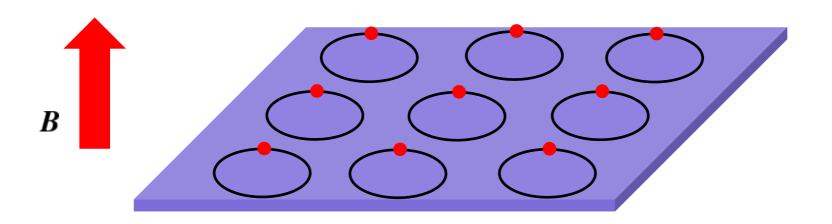
Direct application of Kubo formula in THE often leads to unphysical results:

$$\kappa_{xy}^{\mathrm{Kubo}} \propto \frac{1}{T_0}$$

Electron:
$$\kappa_{xy}^{\text{Kubo}} = \frac{1}{2T_0\hbar V} \sum_{\substack{nk \ 3r}} \text{Im} \left\langle \frac{\partial u_{nk}}{\partial k_x} \middle| \left(\hat{\mathcal{H}}_k + \epsilon_{nk} - 2\mu_0 \right)^2 \middle| \frac{\partial u_{nk}}{\partial k_y} \right\rangle f_{nk}$$
Phonon:
$$\frac{\mathcal{K}_{ubo}}{\mathbb{R}^n} \frac{\hbar}{2V} \sum_{\substack{nk \ 2V = 0 \ k, i=1}}^{\mathbb{R}^n} \frac{\partial \Psi_{ki}}{\partial k_i} \left[\frac{\partial \Psi_{ki}}{\partial k_i} - \frac{\partial \Psi_{ki}}{\partial k_i} \right]$$

L. Zhang *et al.*, Phys. Rev. Lett. **105**, 225901(2010) Katsura, Nagaosa& P. A. Lee, Phys. Rev. Lett. **104**, 066403 (2010)

Circular current component



Electric current: Electromagnetic magnetization

Energy current: Energy Magnetization

$$\frac{\partial \hat{h}(\mathbf{r})}{\partial t}$$
 $\nabla \cdot \hat{\mathbf{r}} \cdot \mathbf{r}$

In equilibrium:

$$\nabla \cdot \boldsymbol{J}_{E}^{\mathrm{eq}}(\boldsymbol{r}) = 0 \qquad \Longrightarrow \qquad \boldsymbol{J}_{E}^{\mathrm{eq}}(\boldsymbol{r}) = \nabla \times \boldsymbol{M}_{E}$$

Transport current

$$\frac{\partial \hat{h}(r)}{\partial t} \nabla \hat{J} \int_{E} \nabla \hat{L} \hat{L}$$

Current is defined only up to a curl:

$$oldsymbol{J}_E$$
 and $oldsymbol{J}_E - oldsymbol{
abla} imes oldsymbol{M}_E$

What are we measuring in transport experiments?

- The curl uncertainty does not affect the total current measured
- We can define a transport current that vanishes when equilibrium:

$$\boldsymbol{J}_E^{\mathrm{tr}} = \boldsymbol{J}_E - \boldsymbol{\nabla} \times \boldsymbol{M}_E$$

Einstein Relations

$$\mu_e = \frac{eD}{k_B T}$$

Transport current vanishes in the equilibrium state

Einstein relations

Electron current
$$J_x = \mu_e nE - D \frac{dn}{dx}$$

Equilibrium state $J_x = 0$

Equilibrium distribution
$$n(x) = N(T) \exp \left\{ -\frac{\epsilon - e\phi(x) - \mu}{k_B T} \right\}$$

Gravitation Field and Thermal Transport

Introducing gravitation field:

$$\hat{H} = \int d\mathbf{r} \hat{h}(\mathbf{r}) \implies \hat{H}^{\psi} = \int d\mathbf{r} \left[1 + \psi(\mathbf{r})\right] \hat{h}(\mathbf{r})$$

Equilibrium state:

$$\rho_{\text{eq}} = \frac{1}{Z} e^{-\int d\mathbf{r} \beta \left[1 + \psi(\mathbf{r})\right] \hat{h}(\mathbf{r})} = \frac{1}{Z} e^{\int d\mathbf{r} \frac{\hat{h}(\mathbf{r})}{k_B T(\mathbf{r})}} \Longrightarrow \beta = \frac{1}{k_B T \left(1 + \psi(\mathbf{r})\right)}$$

$$\beta = \text{Constant}, \qquad \nabla \psi + \frac{\nabla T}{T} = 0$$

Einstein relations $J_E^{\text{tr}} = \tilde{L} \nabla \psi + L_T^{\frac{1}{T}} \nabla T$ $\Longrightarrow \tilde{L} = L, \quad \kappa = \frac{\tilde{L}}{T}$ $J_E^{\text{tr}} \propto \nabla \beta$

J. M. Luttinger, Phys. Rev. **135**, A1505(1964)

Magnetization Correction

$$J_E \to J_E^{\psi} = (1 + \psi)^2 J_E$$

$$M_E \to M_E^{\psi} = (1 + \psi)^2 M_E$$

$$\times M_E (\mu, T) \to \nabla \times M_E^{\psi} = (1 + \psi)^2 \nabla \times M_E + 2\nabla \psi \times M_E \nabla$$

Transport current

$$\boldsymbol{J}_{E}^{\mathrm{tr}} = \boldsymbol{J}_{E} - \boldsymbol{\nabla} \times \boldsymbol{M}_{E}$$

$$\Delta \boldsymbol{J}_{E}^{\text{tr}} = -L\nabla\psi - 2\nabla\psi \times \boldsymbol{M}_{E}, \qquad \kappa_{xy}^{\text{tr}} = \frac{L}{2M_{E}^{2}} + \frac{2M_{E}^{2}}{2M_{E}^{2}}$$

Cooper, Halperin & Ruzin, Phys. Rev. B 55, 2344 (1997)

Remaining Issues

• How to calculate (energy) magnetization(s) in an open system?

Rigorous derivation for the magnetization correction: Einstein relations maintained?

Tao Qin, Qian Niu and Juren Shi, *Energy magnetization and thermal Hall effect*. Phys. Rev. Lett. **107**, 236601(2011)

Theoretical Difficulty

Orbital magnetization:
$$\hat{m{M}} = -\frac{e}{2} m{r} imes \hat{m{v}}$$

OM is simply the equilibrium expectation value of \hat{M}

$$oldsymbol{M} = \left\langle \Psi_G \left| \hat{oldsymbol{M}} \left| \Psi_G
ight
angle
ight.$$

However, for crystalline solid: $\psi_{n\mathbf{k}} = e^{i\mathbf{k}\cdot\mathbf{r}}u_{n\mathbf{k}}(\mathbf{r})$

 $\langle \psi_{n\mathbf{k}} | \hat{\mathbf{M}} | \psi_{n\mathbf{k}} \rangle$ has no deterministic expectation value!

Theory of Orbital Magnetization:

J. Shi, G. Vignale, D. Xiao, Q. Niu, Phys. Rev. Lett. **99**, 197202 (2007). Thonhauser, Ceresoli, Vanderbilt, Resta, PRL **95**, 137205 (2005). D. Xiao, J. Shi and Q. Niu, Phys. Rev. Lett. **95**, 137204 (2005).

However, it is only applicable to electrons in crystalline solids

Magnetization Formulas

$$-\frac{\partial \mathbf{M}_{N}}{\partial \mu_{0}} = \frac{\beta_{0}}{2i} \nabla_{\mathbf{q}} \times \left\langle \hat{n}_{-\mathbf{q}}; \hat{\mathbf{J}}_{N,\mathbf{q}} \right\rangle_{0} \Big|_{\mathbf{q} \to 0}$$

$$\mathbf{M}_{N} - T_{0} \frac{\partial \mathbf{M}_{N}}{\partial T_{0}} = \frac{\beta_{0}}{2i} \nabla_{\mathbf{q}} \times \left\langle \hat{K}_{-\mathbf{q}}; \hat{\mathbf{J}}_{N,\mathbf{q}} \right\rangle_{0} \Big|_{\mathbf{q} \to 0}$$

$$-\frac{\partial \mathbf{M}_{Q}}{\partial \mu_{0}} = \frac{\beta_{0}}{2i} \nabla_{\mathbf{q}} \times \left\langle \hat{n}_{-\mathbf{q}}; \hat{\mathbf{J}}_{Q,\mathbf{q}} \right\rangle_{0} \Big|_{\mathbf{q} \to 0}$$

$$2\mathbf{M}_{Q} - T_{0} \frac{\partial \mathbf{M}_{Q}}{\partial \mathbf{m}_{Q}} = \frac{\beta_{0}}{2i} \nabla_{\mathbf{q}} \times \left\langle \hat{K}_{-\mathbf{q}}; \hat{\mathbf{J}}_{Q,\mathbf{q}} \right\rangle_{0} \Big|_{\mathbf{q} \to 0}$$

Pre-requisite: In the presence of gravitation field ψ and potential ϕ , the current operators should scale with:

$$\hat{\boldsymbol{I}}_{Q}(\boldsymbol{r}) \equiv \hat{\boldsymbol{I}}_{E}(\boldsymbol{r}) - \mu_{0} \hat{\boldsymbol{I}}_{N}(\boldsymbol{r}) \qquad \boldsymbol{M}_{Q} \equiv \boldsymbol{M}_{E} - \mu_{0} \boldsymbol{M}_{N} \qquad \hat{\boldsymbol{K}}(\boldsymbol{r}) \equiv \hat{\boldsymbol{h}}(\boldsymbol{r}) - \mu_{0} \hat{\boldsymbol{n}}(\boldsymbol{r})$$

Magnetization Corrections

$$\begin{bmatrix} \boldsymbol{J}_{1}^{\text{tr}} \\ \boldsymbol{J}_{2}^{\text{tr}} \end{bmatrix} = \begin{bmatrix} \overleftrightarrow{L}^{(11)} & \overleftrightarrow{L}^{(12)} - \frac{\boldsymbol{M}_{N}}{\beta_{0}V} \times \\ \overleftrightarrow{L}^{(21)} - \frac{\boldsymbol{M}_{N}}{\beta_{0}V} \times & \overleftrightarrow{L}^{(22)} - \frac{2\boldsymbol{M}_{Q}}{\beta_{0}V} \times \end{bmatrix} \begin{bmatrix} \boldsymbol{X}_{1} \\ \boldsymbol{X}_{2} \end{bmatrix}$$

	Current J	Force X
1	Particle current	Electric Field, Density gradient
2	Heat Current	Temperature Gradient, Gravitation Field

$$J_1^{\phi,\psi} = J_N^{\phi,\psi} \qquad J_2^{\phi,\psi} = \hat{J}_Q^{\phi,\psi} = J_E^{\phi,\psi} - \alpha(r)J_N^{\phi,\psi}$$

$$\alpha(r) = [1 + \psi(r)][\phi(r) + \mu(r)] \qquad E = I/k_B I + \psi(r) = I/F$$

$$X_1 = -\beta(r)\nabla\alpha(r) \qquad X_2 = \nabla\beta(r)$$

Proof--Canonical formulas



$$\kappa_{11}^{q=0}(\mathbf{r}) = \frac{\partial M_N(\mathbf{r})}{\partial \mu_0} \bigg|_{T_0}, \kappa_{12}^{q=0}(\mathbf{r}) = T_0 \frac{\partial M_N(\mathbf{r})}{\partial T_0} \bigg|_{\mu_0},$$

$$\kappa_{21}^{q=0}(\mathbf{r}) = \frac{\partial M_Q(\mathbf{r})}{\partial \mu_0} \bigg|_{T_0} + M_N(\mathbf{r}), \kappa_{22}^{q=0}(\mathbf{r}) = T_0 \frac{\partial M_Q(\mathbf{r})}{\partial T_0} \bigg|_{\mu_0}.$$

 ∇_{q} both sides of $\chi_{ij}^{q}(r)$, let $q \to 0$ and integrate over r.

Proof—Magnetization corrections

Density matrix
$$\hat{\rho} \approx \hat{\rho}_{leq} + \hat{\rho}_{1}$$

$$\hat{\rho}_{leq} = \frac{1}{Z} \exp \left[-\int d\mathbf{r} \left(\hat{h}(\mathbf{r}) - \mu(\mathbf{r}) \hat{n}(\mathbf{r}) \right) / (k_B T(\mathbf{r})) \right]$$

$$\hat{\rho}_{1} \text{ is determined by } i\hbar \frac{\partial \hat{\rho}}{\partial t} + [\hat{\rho}, \hat{H}_{\phi,\psi}] = 0$$

$$J_i^{\phi,\psi} = J_i^{\text{leq}} + J_i^{\text{Kubo}}, \ J_i^{\text{leq}} = \text{Tr}\hat{\rho}_{\text{leq}}\hat{J}_i^{\phi,\psi}, \ J_i^{\text{Kubo}} = \text{Tr}\hat{\rho}_1\hat{J}_i^{\phi,\psi}$$
Linear order: $\mu(\mathbf{r}) \approx \mu_0 + \delta\mu(\mathbf{r}), \frac{1}{2}\mathcal{F}_{\text{const}}$

Local equilibrium current:

$$J_i^{\text{leq}}(\mathbf{r}) \approx J_i^{\text{eq}}(\mathbf{r}) + \sum_{j=1}^2 \int d\mathbf{r}' \chi_{ij}(\mathbf{r}, \mathbf{r}') \chi_j(\mathbf{r}')$$

$$\chi_1(\mathbf{r}) \equiv \delta \mu(\mathbf{r}), \quad \chi_2(\mathbf{r}) \equiv -T_0 \delta[1/T(\mathbf{r})]$$
Equilibrium current:

$$\boldsymbol{J}_{1}^{\mathrm{eq}}(\boldsymbol{r}) = \begin{bmatrix} 1 + \psi(\boldsymbol{r}) \end{bmatrix} \boldsymbol{\nabla} \times \boldsymbol{M}_{N}(\boldsymbol{r})$$

$$\boldsymbol{J}_{2}^{\text{eq}}(\boldsymbol{r}) = [1 + \psi(\boldsymbol{r})]^{2} [\boldsymbol{\nabla} \times \boldsymbol{M}_{E}(\boldsymbol{r}) - \mu(\boldsymbol{r}) \boldsymbol{\nabla} \times \boldsymbol{M}_{N}(\boldsymbol{r})]$$

Proof—Magnetization corrections

From
$$q = -iq \times M_{ij} = -iq \times M_{ij}$$
, we have:

$$J_{1}^{\text{leq}}(\mathbf{r}) \approx \nabla \times M_{N}^{\phi,\psi}(\mathbf{r}) - \frac{1}{2} M_{N}(\mathbf{r}) \times X_{2}$$

$$= \frac{1}{2} \frac{e^{\phi,\psi}}{2} \left(\frac{1}{2} \right) \times \frac{1}{2} \frac{1}{2}$$

Introducing the transport currents:

$$oldsymbol{J}_{N(E)}^{\phi,\psi, ext{tr}} = oldsymbol{J}_{N(E)}^{\phi,\psi} - oldsymbol{
above} oldsymbol{X}_{N(E)}^{\phi,\psi} = oldsymbol{J}_{i}^{ ext{eq}} + oldsymbol{J}_{i}^{ ext{t}}$$

$$oldsymbol{J}_{i}^{Kubo} \approx \sum_{j} \overleftarrow{L}^{(ij)} \cdot X_{j}.$$

Application to the anomalous Hall system

The electron energy density:

$$\hat{h}(\mathbf{r}) = \left\{ \frac{m}{2} \left[\hat{\mathbf{v}} \hat{\varphi}(\mathbf{r}) \right]^{\dagger} \cdot \left[\hat{\mathbf{v}} \hat{\varphi}(\mathbf{r}) \right] + \hat{\varphi}^{\dagger}(\mathbf{r}) V(\mathbf{r}) \hat{\varphi}(\mathbf{r}) \right\}$$

Energy current operator:

$$\hat{\mathbf{J}}_{E}(\mathbf{r}) = \frac{1}{2} \left\{ \left[\hat{\mathbf{v}} \hat{\varphi}(\mathbf{r}) \right]^{\dagger} \left[\hat{\mathcal{H}} \hat{\varphi}(\mathbf{r}) \right] + h.c. \right\}$$

$$\hat{\mathcal{H}} = \frac{m}{2} \hat{\mathbf{v}}^{2} + V(\mathbf{r})$$

• Gravitational field $\psi \neq 0$, $\hat{h}^{\psi}(\mathbf{r}) = [1 + \psi(\mathbf{r})] \hat{h}(\mathbf{r})$

$$\hat{\boldsymbol{J}}_{E}^{\psi}(\boldsymbol{r}) = [1 + \psi(\boldsymbol{r})]^{2} \hat{\boldsymbol{J}}_{E}(\boldsymbol{r}) + \nabla (1 + \psi(\boldsymbol{r}))^{2} \times \hat{\boldsymbol{\Lambda}}(\boldsymbol{r})$$

$$\hat{\mathbf{\Lambda}}(\mathbf{r}) = \frac{\hbar}{8i} (\hat{\mathbf{v}}\hat{\varphi})^{\dagger} \times (\hat{\mathbf{v}}\hat{\varphi})$$

Gauge freedom--curl:

$$\nabla \times \left((1 + \psi(\mathbf{r}))^2 \hat{\mathbf{\Lambda}}(\mathbf{r}) \right)$$

New current operator:

$$\hat{\boldsymbol{J}}_{E}(\boldsymbol{r}) \rightarrow \hat{\boldsymbol{J}}_{E}\left(\boldsymbol{r}\right) - \boldsymbol{\nabla} \times \hat{\boldsymbol{\Lambda}}\left(\boldsymbol{r}\right)$$

Application to the anomalous Hall system

Kubo formula:

$$\kappa_{xy}^{\text{Kubo}} = \frac{1}{2T_0\hbar V} \sum_{nk} \text{Im} \left(\frac{\partial u_{nk}}{\partial k_x} \right) \left(\hat{\mathcal{H}}_k + \epsilon_{nk} - 2\mu_0 \right)^2 \left[\frac{\partial u_{nk}}{\partial k_x} \right] f_{nk}$$

Energy magnetization:

$$J_{Q,q} = \frac{\partial M_Q}{\partial T_0} = \frac{\partial M_Q}{\partial T_0} = \frac{\partial M_Q}{\partial T_0} = \frac{\partial M_Q}{\partial T_0} \times \left(K - \frac{\tilde{q}}{q}\right)$$

$$\tilde{h} = \frac{1}{\sqrt{\partial k_x}} \operatorname{Im} \left[\frac{\partial u_{nk}}{\partial k_y} \right] (H_{k_y} - 2u_{nk})^2 \left[\frac{\partial u_{nk}}{\partial k_y} \right] f.$$

$$\sum_{nk} \operatorname{Im} \left[\frac{\partial u_{nk}}{\partial k_{x}} \Big|_{(\mathcal{E}_{nk} - \mathcal{H}_{k})^{2} - 4(\mathcal{E}_{nk} - \mu_{0})} (\mathcal{E}_{nk} - \mathcal{H}_{k})} \frac{\partial u_{nk}}{\partial k_{y}} \right] \qquad \frac{1}{4\hbar}$$

$$\times (\epsilon_{nk} - \mu_{0}) J_{nk}$$

Application to the anomalous Hall system

$$\kappa_{xy}^{\text{tr}} \equiv \kappa_{xy}^{\text{Kubo}} + \frac{2M_Q^z}{T_0 V}$$

$$\kappa_{xy}^{\text{tr}} = -\frac{1}{e^2 T_0} \int d\epsilon (\epsilon - \mu_0)^2 \sigma_{xy}(\epsilon) \frac{df(\epsilon)}{d\epsilon}$$

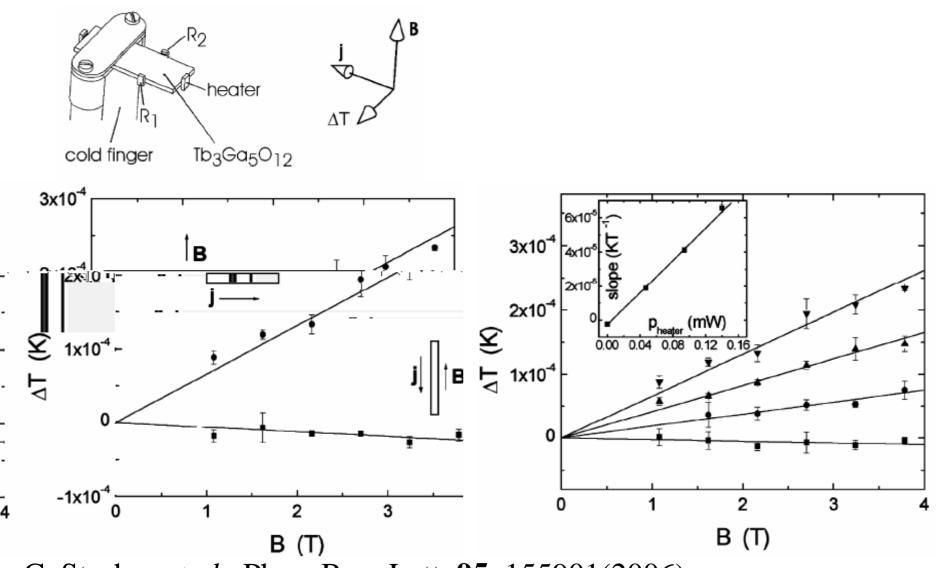
$$\sigma_{xy}(\epsilon) = -\frac{e^2}{\hbar} \sum_{\epsilon_{nk} \le \epsilon} \Omega_{nk}^z \qquad \Omega_{nk}^z \equiv -2 \operatorname{Im} \left\langle \frac{\partial u_{nk}}{\partial k_x} \middle| \frac{\partial u_{nk}}{\partial k_y} \right\rangle$$

Wiedemann-Franz law:

Theory for the phonon Hall effect

- Experiments on phonon Hall effect
- Issues of existing theories
- Our theory
 - General phonon dynamics for magnetic systems
 - Proper evaluation of phonon Hall coefficient
 - Topological phonon system
 - Low temperature behavior
 - Tao Qin and Junren Shi, *Berry curvature and phonon Hall effect*, arXiv: 1111. 1322v1

Experiments on phonon Hall effect

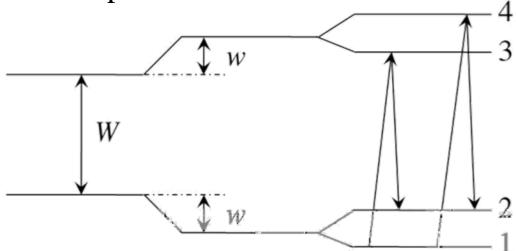


C. Strohm et al., Phys. Rev. Lett. 95, 155901(2006)

A. Inyushkin et al., JETP Lett. 86, 379 (2007)

Existing theories: Spin-lattice Raman interaction

Microscopic model:



$$H_R = K \sum_m \mathbf{M} \cdot \Omega_m$$
$$\Omega_m = \mathbf{u}_m \times \mathbf{p}_m$$

Kronig, Physica **6**, 33(1939)

L. Sheng et al., Phys. Rev. Lett. 96, 155901(2006)

Yu. Kagan et al., Phys. Rev. Lett. 100, 145902 (2008)

L. Zhang, et al., Phys. Rev. Lett. 105, 225901(2010)

Kubo formula or its equivalents is employed

Issue #1: inappropriate microscopic model

$$H_R = K \sum_m \mathbf{M} \cdot \Omega_m$$
$$\Omega_m = \mathbf{u}_m \times \mathbf{p}_m$$

Considering a rigid-body motion: $u_m = u$

$$H_R \to K \boldsymbol{M} \cdot (\boldsymbol{u} \times \boldsymbol{P})$$

A magnetic solid will experience a Lorentz force!?

The microscopic model breaks Principle of Relativity!

Issue #2: Kubo formula

$$\kappa_{xy}^{\text{Kubo}} = \frac{1}{V k_B T^2} \lim_{s \to 0} \lim_{q \to 0} \int_0^\infty dt e^{-st} \left\langle \hat{J}_{E,-q}^y; \, \hat{J}_{E,q}^x(t) \right\rangle$$

However, this formula is not applicable for magnetic systems!

$$\kappa_{xy}^{\text{tr}} \equiv \kappa_{xy}^{\text{Kubo}} + \frac{2M_{Q}^{2}}{T_{0}V}$$

$$2M_{Q} - T_{0}\frac{\partial M_{Q}}{\partial T_{0}} = \frac{\beta_{0}}{2i} \nabla_{q} \times \left\langle \hat{K}_{-q}; \hat{J}_{Q,q} \right\rangle_{0} \Big|_{q \to 0}$$

Tao Qin, Qian Niu and Juren Shi, Energy magnetization and thermal Hall effect. Phys. Rev. Lett. **107**, 236601(2011)

Our theory: Phonon Dynamics

- The electron Berry phase —— The effective magnetic field
- The effective Hamiltonian

$$\frac{1}{2M_{\kappa}} \frac{(-i\hbar \nabla_{l\nu} - A_{l\nu} (\{R\}))^{2}}{2M_{\kappa}} + V_{\text{eff}}(R)H = \sum_{l\kappa} A_{l\kappa} (\{R\}) \equiv i\hbar \langle \Phi_{0} (\{R\}) | \nabla_{l\kappa} \Phi_{0} (\{R\}) \rangle$$

Mead-Truhlar term:
$$A_{l\kappa}(\{R\}) \cdot \frac{\hbar}{i} \nabla_{l\kappa}$$

C. A. Mead and D. G. Truhlar, J. Chem. Phys. **70**, 2284 (1979)

Effective magnetic field acting on phonons

The effective magnetic field:

$$G_{\alpha\beta}^{\kappa\kappa'}(\boldsymbol{R}_{l}^{0}-\boldsymbol{R}_{l'}^{0})=2\hbar\mathrm{Im}\left\langle\frac{\partial\Phi_{0}}{\partial u_{\beta,l'\kappa'}}\bigg|\frac{\partial\Phi_{0}}{\partial u_{\alpha,l\kappa}}\right\rangle\bigg|_{u_{l\kappa}\to0}$$

A constraint naturally emerges from the translational symmetry:

$$\sum_{l\kappa\kappa'} G_{\alpha\beta}^{\kappa\kappa'}(\boldsymbol{R}_l^0) = 0$$

Principle of Relativity recovers.

Phonon dynamics and Berry curvature

The equations of motion

$$\begin{split} \dot{\tilde{u}}_k &= P_k \\ \dot{P}_k &= -D_k \tilde{u}_k + G_k P_k \\ \omega_{ki} \Psi_{ki} &= \begin{pmatrix} 0 & \mathrm{i} \\ -\mathrm{i} D_k & \mathrm{i} G_k \end{pmatrix} \Psi_{ki} \equiv \tilde{H}_k \Psi_{ki} \end{split}$$

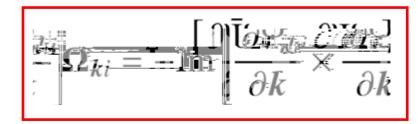
6r branches of phonons satisfing:

$$\omega_{ki}^{(-)} = -\omega_{-ki}^{(+)} \quad \Psi_{ki}^{(-)} = \Psi_{-ki}^{(+)*}$$

$$\bar{\Psi}_{ki}\Psi_{ki} = 1 \qquad \bar{\Psi}_{ki} = \Psi_{ki}^{\dagger} \tilde{D}_{k}$$

The phonon Berry connection and Berry curvature

$$\mathcal{A}_{ki} = i\bar{\Psi}_{ki} \frac{\partial \Psi_{ki}}{\partial k}$$



Phonon Hall coefficient

Kubo formula

$$\mathcal{L}_{ki}^{\text{Pub}} \frac{\hbar}{\omega_{ki}} \frac{3r}{n_{ki} + \frac{3r}{2}} \mathcal{L}_{xy}^{\text{Ruso}} = \frac{1}{VT_0} \sum_{k:i=1}^{1} \mathcal{L}_{xy}^{\text{Ruso}}$$

$$\mathcal{M}_{ki} = \operatorname{Im} \left[\frac{\partial \bar{\psi}_{ki}}{\partial k} \times \tilde{H}_k \frac{\partial \psi_{ki}}{\partial k} \right]$$

Energy magnetization

$$\tilde{M}_{E}^{z} = -\frac{\hbar}{2} \sum_{k;i=1}^{3r} \left[\Omega_{i}^{z} \omega_{i}^{3} n_{i}' + M_{i}^{z} (2\omega_{ki}n_{ki} + \omega_{i}^{2} n_{i}' + 1) \right]$$

$$2M_E^z - T\frac{\partial M_E^z}{\partial T} = \tilde{M}_E^z$$

Phonon Hall coefficient

$$\kappa_{xy}^{\text{tr}} \equiv \kappa_{xy}^{\text{Kubo}} + \frac{2M_Q^2}{T_0 V}$$

$$\kappa_{xy}^{\text{tr}} = -\frac{(\pi k_B)^2}{3h} Z_{\text{ph}} T - \frac{1}{T} \int d\epsilon \epsilon^2 \sigma_{xy}(\epsilon) \frac{dn(\epsilon)}{d\epsilon}$$

$$Z_{\rm ph} = \frac{2\pi}{V} \sum_{k:i=1}^{3r} \Omega_{ki}^{z}, \qquad \sigma_{xy}(\epsilon) = -\frac{1}{V\hbar} \sum_{\hbar\omega_{ki} \le \epsilon} \Omega_{ki}^{z}$$

Topological Phonon System

$$Z_{\rm ph} \neq 0$$

$$\kappa_{xy}^{\text{topo.}} = -\frac{(\pi k_B)^2}{3h} Z_{\text{ph}} T$$

$$Z_{\rm ph} = \begin{cases} \text{Integer,} & 2D\\ \frac{G_z}{2\pi}, & 3D \end{cases}$$

 G_z : z-component of a reciprocal lattice vector G

Halperin, Jpn. J. Appl. Phys. 26S3, 1913 (1987)

Our theory: long wave limit

Constraint on the effective magnetic field acting on atoms:

$$\sum_{l \nu \nu'} G_{\alpha\beta}^{\kappa\kappa'}(\boldsymbol{R}_l^0) = 0$$

Phonon Hall coefficient

$$\kappa_{xy}^{\rm tr} \propto T^3$$

Instead of $\kappa_{xy}^{\text{tr}} \propto T$

L. Sheng *et al.*, Phys. Rev. Lett. **96**, 155901(2006). J. Wang *et al.*, Phys. Rev. B **80**, 012301 (2009)

Summary

$$-\frac{\partial \boldsymbol{M}_{N}}{\partial \mu_{0}} = \frac{\beta_{0}}{2i} \boldsymbol{\nabla}_{\boldsymbol{q}} \times \left\langle \hat{\boldsymbol{n}}_{-\boldsymbol{q}}; \hat{\boldsymbol{J}}_{N,\boldsymbol{q}} \right\rangle_{0} \Big|_{\boldsymbol{q} \to 0}$$

$$\boldsymbol{M}_{N} - T_{0} \frac{\partial \boldsymbol{M}_{N}}{\partial T_{0}} = \frac{\beta_{0}}{2i} \boldsymbol{\nabla}_{\boldsymbol{q}} \times \left\langle \hat{\boldsymbol{K}}_{-\boldsymbol{q}}; \hat{\boldsymbol{J}}_{N,\boldsymbol{q}} \right\rangle_{0} \Big|_{\boldsymbol{q} \to 0}$$

$$-\frac{\partial \boldsymbol{M}_{Q}}{\partial \mu_{0}} = \frac{\beta_{0}}{2i} \boldsymbol{\nabla}_{\boldsymbol{q}} \times \left\langle \hat{\boldsymbol{n}}_{-\boldsymbol{q}}; \hat{\boldsymbol{J}}_{Q,\boldsymbol{q}} \right\rangle_{0} \Big|_{\boldsymbol{q} \to 0}$$

$$2\boldsymbol{M}_{Q} - T_{0} \frac{\partial \boldsymbol{M}_{Q}}{\partial T_{0}} = \frac{\beta_{0}}{2i} \boldsymbol{\nabla}_{\boldsymbol{q}} \times \left\langle \hat{\boldsymbol{K}}_{-\boldsymbol{q}}; \hat{\boldsymbol{J}}_{Q,\boldsymbol{q}} \right\rangle_{0} \Big|_{\boldsymbol{q} \to 0}$$

Summary

- A general phonon dynamics for magnetic systems
- Emergent "magnetic field" for phonon
- Phonon Hall coefficient and phonon Berry curvature
- Topological phonon systems Quantum Hall Effect of Phonon Systems
- Low temperature behavior of ordinary phonon systems: T^3 instead of T
- Linear *T* with quantized coefficient may suggest Topological Phonon System

Tao Qin, Qian Niu and Juren Shi, *Energy magnetization and thermal Hall effect* Phys. Rev. Lett. **107**, 236601(2011)

Tao Qin and Junren Shi, *Berry curvature and phonon Hall effect*, arXiv: 1111. 1322